
Associated Graviton production in Bhabha Scattering[¶]

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With

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⊙ Theories with extra space like dimension(s) have got lots of attention as a solution to the Naturalness problem

⊙ Two main types of theories :

i. Models of large extra dimensions (ADD).*

ii. Models of warped space-time (RS).†

*Arkani-Hamed, Dimopoulos, Dvali [PLB429:263-272,98]

†Randall, Sundrum [PRL83:3370-3373,99]

ADD scenario

- ⊙ $(4 + \delta)$ dimensional theory with extra δ dimension(s) which are compactified (with radius R).
- ⊙ Standard Model fields are confined on $(1+3)$ -dimensional brane though gravity can propagate anywhere in the $(4 + \delta)$ dimensional bulk.
- ⊙ Our 4 dimensional Planck scale (M_p) is related to the $(4 + \delta)$ dimensional Planck scale (\bar{M}_s) which is actually fundamental scale in this scenario.

$$M_p^2 = 8\pi \cdot V_\delta \bar{M}_s^{\delta+2} R^\delta$$

Volume element $V_\delta = (2\pi)^\delta$ assuming toroidal compactification of extra dimensions.

- ⊙ By choosing large R (exp. limit $\approx \# \text{ mm}$), fundamental scale ($\bar{M}_s \sim 1 \text{ TeV}$) can produce the Planck scale ($M_p \sim 10^{19} \text{ GeV}$) in 4 dimension.

⊙ δ compact extra spatial dimensions, \Rightarrow Infinite tower of Kaluza-Klein states with masses

$$m_n^2 \sim \frac{\vec{n}^2}{R^2} \quad \vec{n} = (n_1, n_2, \dots, n_\delta)$$
$$n_i = 0, \pm 1, \pm 2, \dots$$

⊙ The coupling of each graviton KK states to the SM fields remain small, being proportional to $1/M_p$.

⊙ But cumulative effect from full tower of KK states, considering KK state density

$$\rho(m)dm = \left[\frac{2\pi^{\frac{\delta}{2}}}{\Gamma(\frac{\delta}{2})} \right] \frac{\bar{M}_p^2}{M_s^{2+\delta}} m^{\delta-1} dm$$

\longrightarrow keeps the cross-section sizable to have collider signature.

e.g. final cross section $\sim \frac{1}{M_s^2}$ for graviton emission (say).

ADD phenomenology

Two types of large extra dimension signals :

- i. Virtual graviton exchange :
⇒ affecting the S.M. scattering signals.
- ii. Real graviton emission :
⇒ missing energy-momentum.

We consider the second case.

A widely studied channel for discovering ADD ‡

$$e^+e^- \rightarrow \gamma G$$

Note: γ is not monochromatic.

One can probe M_s upto §($\sqrt{s} = 1\text{TeV}$, $\mathcal{L} = 200\text{fb}^{-1}$)

unpolarized	90% polarized	no of ex-dim(δ)
4.1 TeV	5.7 TeV	for $\delta = 2$
3.1 TeV	4.0 TeV	for $\delta = 3$
2.5 TeV	3.0 TeV	for $\delta = 4$
2.0 TeV	2.4 TeV	for $\delta = 5$

‡Mirabelli, Perelstein, Peskin [PRL82:2236-2239,99]

§Giudice, Rattazzi, Wells [NPB544:3-38,99]

Alternative channels are always helpful in establishing any new signal as of ADD origin.

Thus we look for,

$$e^+e^- \rightarrow e^+e^-G$$

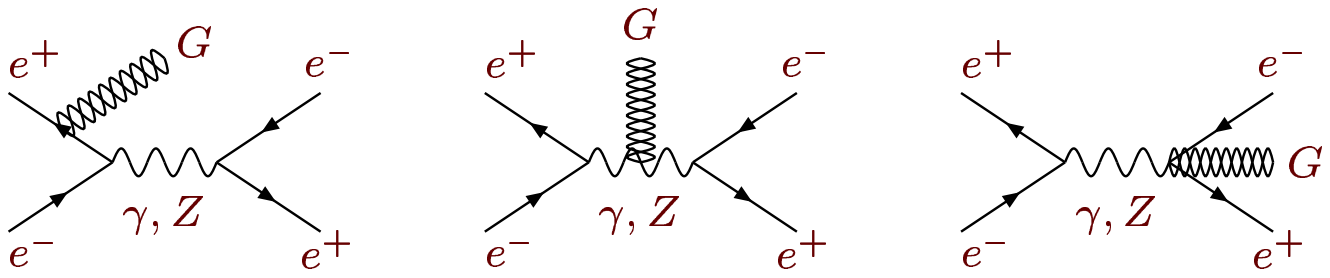
⊙ Bhaba scattering is widely used for calibration of beam.

⊙ So, such scattering accompanied with missing energy (E) can hopefully be measured with precision.

⊙ As, both s and t channel processes are responsible, it may be advantageous over $e^+e^- \rightarrow \mu^+\mu^-G$ ^b due to s channel suppression.

^b *Eboli et al.* [PRD64:035005,01]

⊙ We want to see how far $e^+e^- \rightarrow e^+e^-G$ can probe, side by side with $e^+e^- \rightarrow \gamma G$.



Generic diagrams for the process

$$e^+e^- \rightarrow e^+e^-G$$

- ⊙ As Graviton couples to all particles with energy-momentum, there are **28** Feynman diagrams contributing to the processes both in s & t channel.
- ⊙ **8** more diagrams would be added, if we consider virtual graviton exchange also. But their contributions are insignificantly small.
- ⊙ Full calculation has been done numerically by using the helicity amplitude method. We developed some new **HELAS** routines, which includes **G-F-F**, **G-V-V**, **G-F-F-V** vertices.

Checks :

- ⊙ Calculation of 28 Feynman graphs, even using HELAS, is long and cumbersome, some checks on the numerical results are called for.
- ⊙ Ward identities arising from general coordinate invariance — an essential feature of any theory involving gravity — turn out to be the most useful : with $\epsilon_{\mu\nu}^{(n)}(p_1)$ as the polarisation tensor,

$$M_n(\lambda_1, \lambda_2) = T^{\mu\nu}(\lambda_1, \lambda_2) \epsilon_{\mu\nu}^{(n)*}(p_1)$$

must now satisfy the Ward identities

$$p_1^\mu T_{\mu\nu}(\lambda_1, \lambda_2) = p_1^\nu T_{\mu\nu}(\lambda_1, \lambda_2) = 0$$

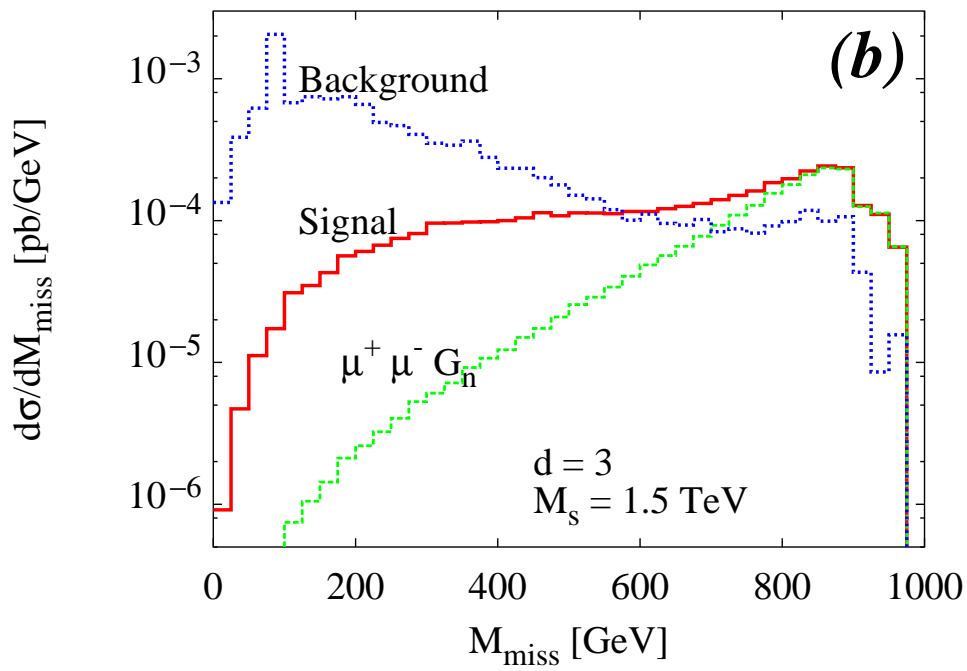
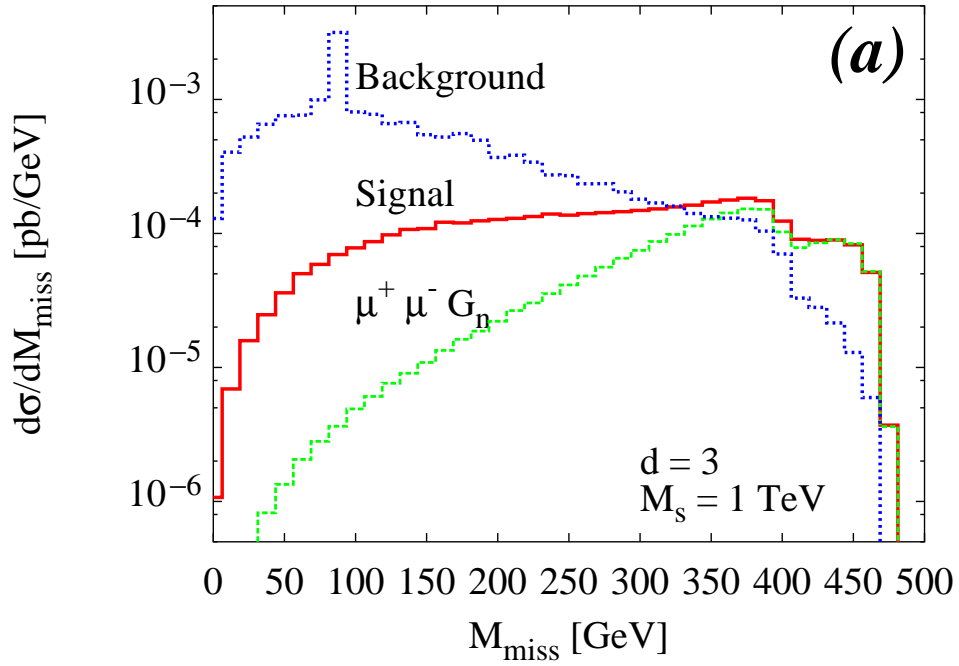
- ⊙ Only the s -channel contributions of our numerical results are in close agreement with $e^+e^- \rightarrow \mu^+\mu^-G$.
- ⊙ So, t -channel contributions can be easily checked using crossing symmetry.

Background :

- ⊙ $e^+e^- \rightarrow e^+e^-\nu_l\bar{\nu}_l$
- ⊙ $e^+e^- \rightarrow e^\pm\nu W^\mp \rightarrow$ Large.
- ⊙ $e^+e^- \rightarrow e^+e^-Z \rightarrow$ Eliminated by missing mass near m_Z

event selection criteria:

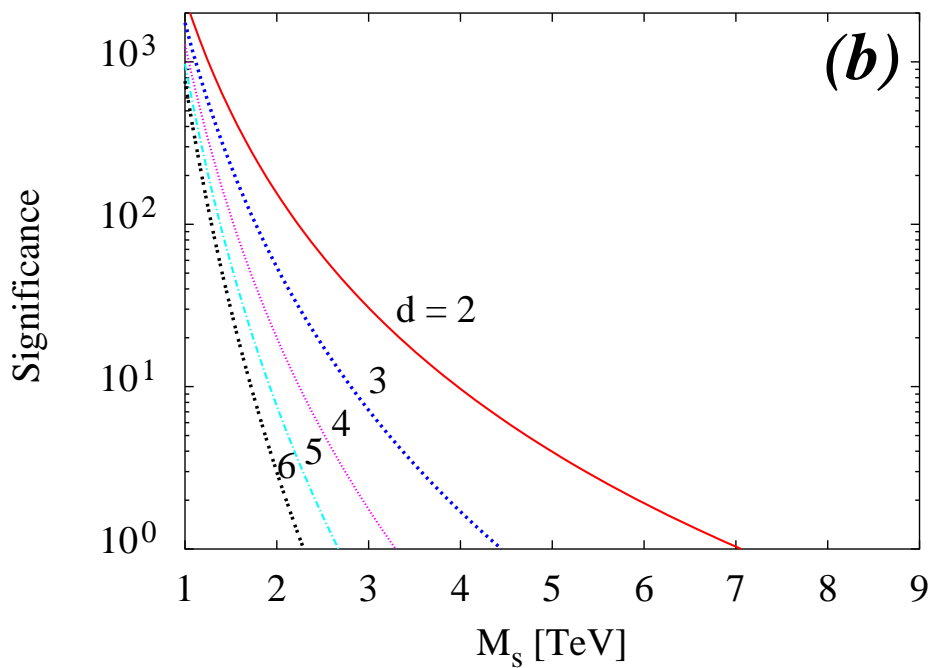
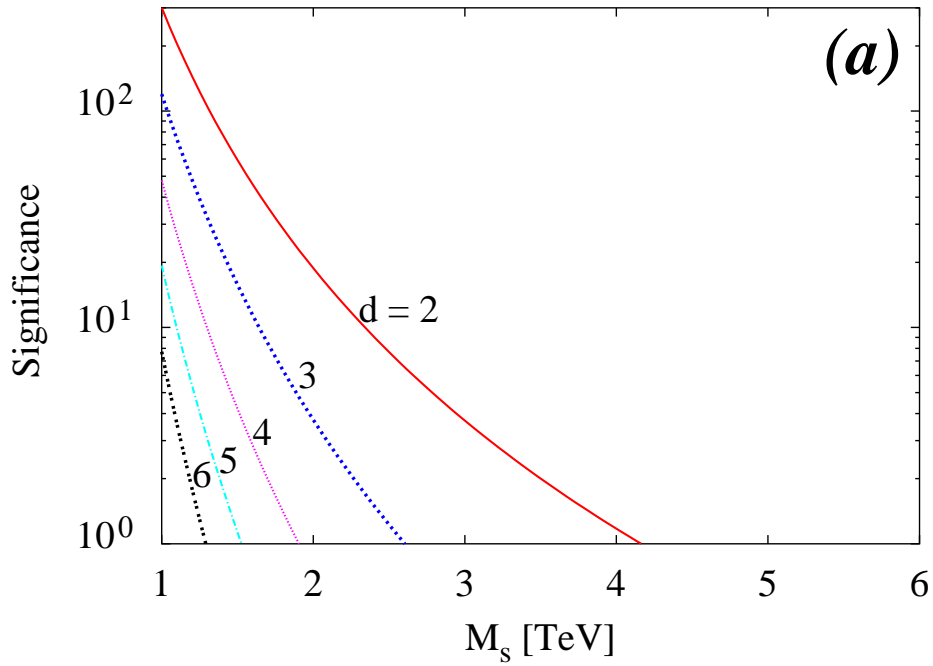
- Electrons are 15° away from the beam pipe
- $(P_T)_e > 10$ GeV
- $P_T(miss) > (15) 25$ GeV for $\sqrt{s} = (500 \text{ GeV}) 1 \text{ TeV}$.
- $\Delta R_{e^+e^-} > 0.2$
- Angle between two electrons $< 175^\circ$
- M_{miss} not inside $m_Z \pm 10$ GeV



⊙ *Missing invariant mass (M_{miss}) distribution for $\sqrt{s} =$
(a) 500 GeV & (b) 1 TeV.*

Missing invariant mass distribution:

- ⊙ Background distribution is characterised by the neutrino pair (missing energy) arising from a real Z .
- ⊙ Missing invariant mass of the signal is much harder than that of the background.
- ⊙ The missing mass is a measure of the energy of the emitted graviton. Now it is well-known that the higher the energy, the more strongly the graviton is coupled to matter and consequently the higher the production cross-section. Moreover, a higher energy leads to a higher density of states, and this further enhances the cross-section.
- ⊙ Peaking behaviour is from radiative return to the Z -pole through graviton emission.
- ⊙ We can therefore obtain a clear separation of signal from background by imposing a cut :
$$M_{\text{miss}} > \begin{cases} (350 \text{ GeV}) & 450 \text{ GeV} \\ (500 \text{ GeV}) & 1 \text{ TeV} \end{cases} \text{ for } \sqrt{s} =$$



⊙ *Significance* $[\sigma^{signal}\sqrt{\mathcal{L}}/\sqrt{\sigma^{bg}}]$ Vs. M_s , for unpolarized beams and $\sqrt{s} =$ (a) 500 GeV (b) 1 TeV.

Used integrated luminosity $\mathcal{L} = 500 fb^{-1}$

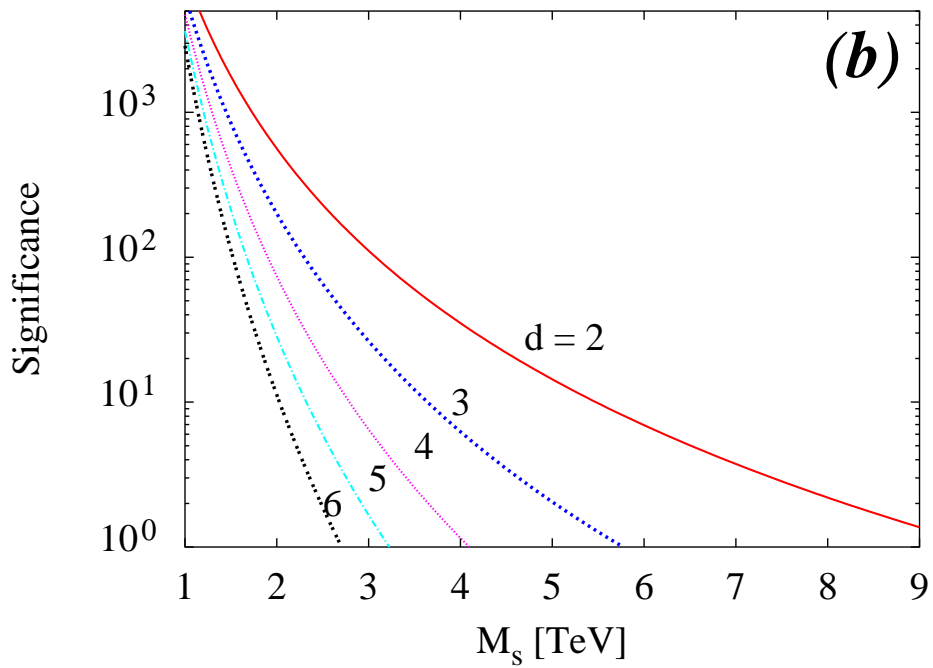
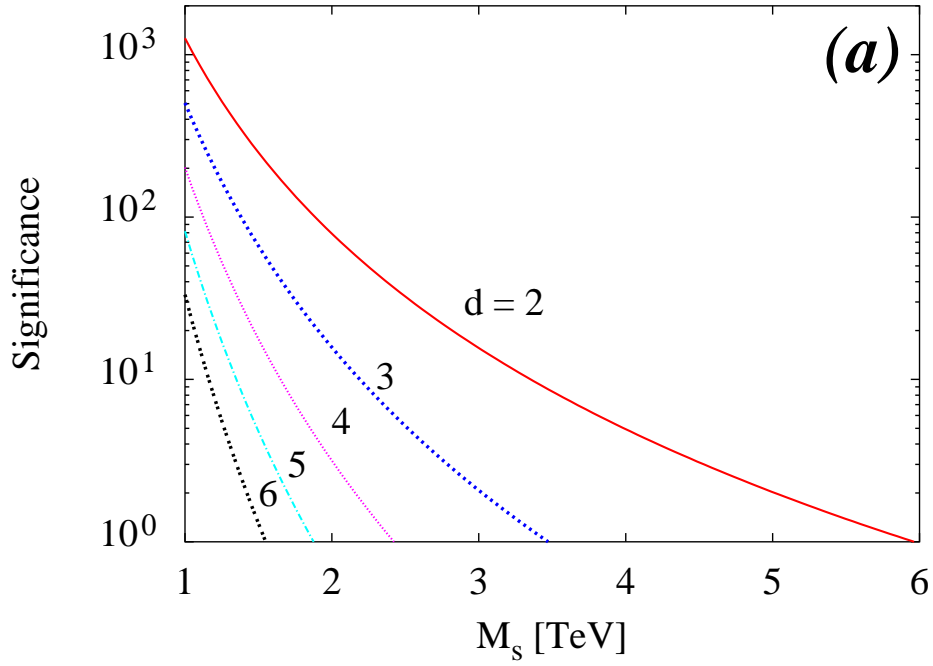
additional criterion $M_{miss} >$ (a) 350 GeV (b) 450 GeV

⊙ Main background : $e^+e^- \rightarrow e^\pm\nu W^\mp$ with further decay of W^\mp

⇒ eliminated by considering polarized beam.

Of course we will lose some signal in this way. But, the experiment with Right-handed e^- and Left-handed e^+ beam would explore farther beyond the reach using unpolarized beam.

⊙ Typical values for the polarization efficiencies used in this analysis are $\mathcal{P}_{e^-} = 0.8$ and $\mathcal{P}_{e^+} = 0.6$.



⊙ *Significance Vs. M_s for polarized beams (80% & 60%) for $\sqrt{s} =$ (a) 500 GeV (b) 1 TeV.*

Used integrated luminosity $\mathcal{L} = 500 fb^{-1}$

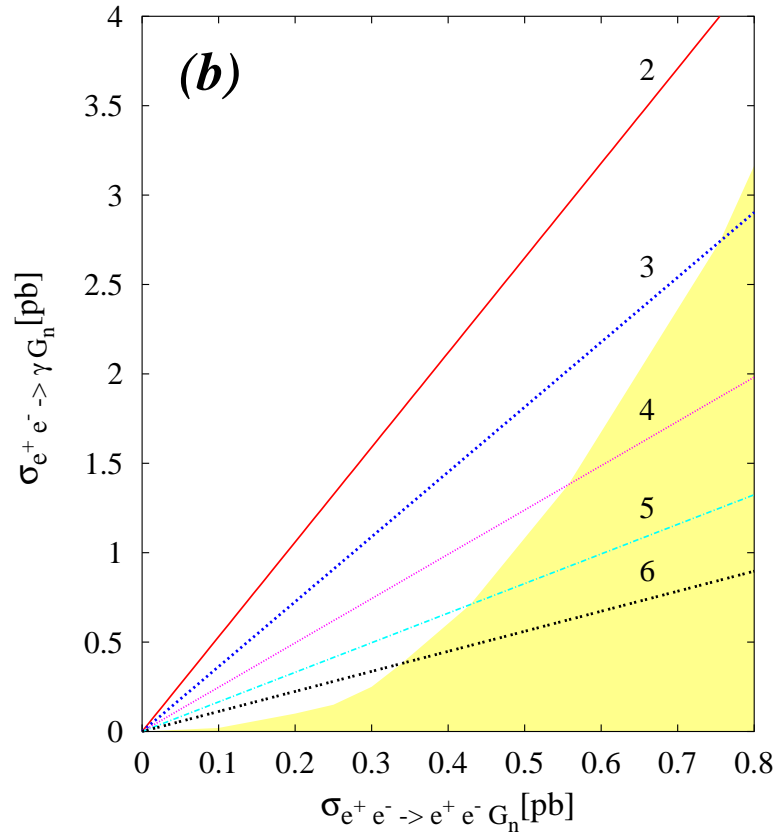
additional criterion $M_{miss} >$ (a) 350 GeV (b) 450 GeV

Results :

- ⊙ Here we have presented some results on string scale (M_s) sensitivity which can be achieved by studying $e^+e^- \rightarrow e^+e^-G$ for various numbers of extra dimensions ($\delta = 2, 3, 4, 5$).
- ⊙ Both Unpolarized and Polarized case has been studied.

Defining the statistical significant discovery to be Significance $[\sigma^{signal}\sqrt{\mathcal{L}}/\sqrt{\sigma^{bg}}] > 3$ at an e^+e^- collider with $\sqrt{s} = 1TeV$ and integrated luminosity $\mathcal{L} = 500fb^{-1}$:

unpolarised	Polarised beams	no of ex-dim(δ)
5.4 TeV	7.4 TeV	for $\delta = 2$
3.6 TeV	4.6 TeV	for $\delta = 3$
2.7 TeV	3.4 TeV	for $\delta = 4$
2.3 TeV	2.8 TeV	for $\delta = 5$
2.0 TeV	2.4 TeV	for $\delta = 6$



⊙ *Correlation plot showing the polarized cross-sections for the processes $e^+e^- \rightarrow \gamma G_n$ and $e^+e^- \rightarrow e^+e^- G_n$ for different values of d . Each curve is generated by varying M_S for $\sqrt{s} = 1$ TeV.*

⊙ A more elegant way of distinguishing these signals from other forms of new physics is to compare the cross-sections arising from *both* the processes $e^+e^- \rightarrow \gamma G_n$ and $e^+e^- \rightarrow e^+e^- G_n$, both of which are determined by the two parameters d and M_S , and hence, will have some correlation.

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- ⊙ Each curve corresponds to variation of M_S over the range allowed by current experimental constraints.
 - ⊙ Using experimental data from a high energy e^+e^- machine, we should be able to pinpoint a small region (ideally a point) on the graphs.
 - ⊙ The position of these would immediately direct one to the value of d ; comparison of the cross-sections would now yield a measurement of the value of M_S .

Distinguishing graviton signal :

- ⊙ Supersymmetric signals (through a pair of sparticle production) can be distinguished by looking at the angular correlation between two leptons.
- ⊙ A more elegant way is to compare the signals arising from *both* the processes $e^+e^- \rightarrow \gamma G_n$ and $e^+e^- \rightarrow e^+e^- G_n \Rightarrow$ possible to pinpoint the value of d as well as M_S .

Conclusions :

- ⊙ To explore large extra dimensions, $e^+e^- \rightarrow \gamma G$ is most sensitive channel, but is not a unique demonstration of ADD.
- ⊙ If $e^+e^- \rightarrow e^+e^- G$ works upto string scale M_s , with a given \sqrt{s} , then one can have a clear idea on the string scale which can be probed in terms of both signals. This will consolidate the search strategy.